# COMP 212 Spring 2023 Homework 08

In this homework you will implement n-body simulation using the Barnes-Hut algorithm. See the instructions at the end of this document for how to compile and run your code: because the project consists of many files, you will need to use CM.make rather than use.

# 1 Sequence Library

For this assignment, you will use the implementation of sequences that you downloaded in lab. It is important that you unzip the homework code so the src directory from sequences and hw08-handout are next to each other. You can find the signature for sequences in the files src/sequence/sequencecore-sig.sml and src/sequence/sequence-sig.sml. For your reference, we describe the functions you can use here:

- Seq.length: 'a Seq.seq -> int Seq.length s evaluates to the number of items in s.
- Seq.empty: unit -> 'a Seq.seq Seq.empty() evaluates to the sequence of length zero.
- Seq.cons: 'a \* 'a Seq.seq -> 'a Seq.seq
  If the length of xs is 1, Seq.cons (x, xs) evaluates to a sequence of length 1+1 whose first item is x and whose remaining 1 items are exactly the sequence xs.
- Seq.singleton: 'a -> 'a Seq.seq
  Seq.singleton x evaluates to a sequence of length 1 where the only item is x.
- Seq.append: 'a Seq.seq -> 'a Seq.seq -> 'a Seq.seq If s1 has length  $l_1$  and s2 has length  $l_2$ , Seq.append evaluates to a sequence with length  $l_1 + l_2$  whose first  $l_1$  items are the sequence s1 and whose last  $l_2$  items are the sequence s2.
- Seq.tabulate: (int -> 'a) \* int -> 'a Seq.seq
  Seq.tabulate (f, n) evaluates to a sequence s with length n where the i<sup>th</sup> item of s is the result of evaluating (f i). Seq.tabulate (f, i) raises Range if n is less than zero.

- Seq.nth: int \* 'a Seq.seq -> 'a

  nth i s evaluates to the *i*<sup>th</sup> item in s. This is zero-indexed. Seq.nth (i, s) will
  raise Range if i is negative or greater than (Seq.length s)-1.
- Seq.filter: ('a -> bool) \* 'a Seq.seq -> 'a Seq.seq
  Seq.filter (p, s) returns the longest subsequence ss of s such that p evaluates to
  true for every item in ss.<sup>1</sup>
- Seq.map: ('a -> 'b) \* 'a Seq.seq -> 'b Seq.seq
  Seq.map (f, s) maps f over the sequence s. That is to say, it evaluates to a sequence s' such that s and s' have the same length and the  $i^{th}$  item in s' is the result of applying f to the  $i^{th}$  item of s.
- Seq.reduce: (('a \* 'a) -> 'a) \* 'a \* 'a Seq.seq -> 'a Seq.reduce c b s combines all of the items in s pairwise with c using b as the base case. c must be associative, with b as its identity.
- Seq.mapreduce : ('a -> 'b) \* 'b \* ('b \* 'b -> 'b) \* 'a Seq.seq -> 'b Seq.mapreduce l e n s is equivalent to Seq.reduce n e (Seq.map l s).
- Seq.toString: ('a -> string) \* 'a Seq.seq -> string Seq.toString (ts, s) evaluates to a string representation of s by using ts to convert each item in s to a string.
- Seq.repeat : int \* a -> 'a Seq.seq
  Seq.repeat (n, x) evaluates to a sequence consisting of exactly n-many copies of x.
- Seq.flatten : 'a Seq.seq Seq.seq -> 'a Seq.seq Seq.flatten ss is equivalent to reduce append (empty ()) ss
- Seq.zip : ('a Seq.seq \* 'b Seq.seq) -> ('a \* 'b) Seq.seq Seq.zip (s1,s2) evaluates to a sequence whose  $n^{th}$  item is the pair of the  $n^{th}$  item of s1 and the  $n^{th}$  item of s2.
- Seq.split: int \* 'a Seq.seq -> 'a Seq.seq \* 'a Seq.seq

  If s has at least i elements, Seq.split (i, s) evaluates to a pair of sequences
  (s1,s2) where s1 has length i and Seq.append (s1, s2) is the same as s. Otherwise it raises Range.
- Seq.take : int \* 'a Seq.seq -> 'a Seq.seq
   Seq.take (i, s) evaluates to the sequence containing exactly the first i elements of s if 0 < i < length s, and raises Range otherwise.</li>
- Seq.drop: int \* 'a Seq.seq -> 'a Seq.seq Seq.drop (i, s) evaluates to the sequence containing all but the first i elements of s if  $0 \le i \le length s$ , and raises Range otherwise.

<sup>&</sup>lt;sup>1</sup>Here we use the term "subsequence" to mean any subsequence of a sequence, not necessecarily one whose elements are consecutive in the original sequence. For example,  $\langle \rangle$ ,  $\langle 3 \rangle$ , and  $\langle 2, 4 \rangle$  are subsequences of  $\langle 1, 2, 3, 4 \rangle$ .

# 2 *n*-Body Simulations

The main portion of this programming assignment is modeling movements of bodies through a universe represented by a two-dimensional Euclidian plane. To make this model, we must pick an SML representation of points in the plane that allows us to meaningfully measure the distance between points—that is to say, we must pick a way to measure the universe. We will represent a point in the plane by a pair of floating point numbers, represented by a pair of values of type real. The type Scalar.scalar is a synonym for real (mathematically, a "scalar" is part of the structure of a "vector space", but we won't go into more detail here).

#### 2.1 The Plane

#### 2.1.1 Points and Vectors

We have provided an implementation of the plane, which consists of points and vectors, based on the code from class. These are represent by the types Plane.point and Plane.vec. In order to write our implementation of the Barnes-Hut algorithm, we need several operations on vectors and points in space, many of which we discussed in lecture. The type Plane.point is used to represent a point in space, and the type Plane.vec is used to represent vectors of velocity, acceleration, etc. In the implementation, we define the type of points and vectors as in lecture:

```
type Plane.point = Scalar.scalar * Scalar.scalar
type Plane.vec = Scalar.scalar * Scalar.scalar
```

In your implementation, these types are *abstract*, which means you should code only in terms of the provided operations on points and vectors.

You can see the full signature for the plane in space.sig. Some operations include:

- Plane.vecFromPoints: Plane.point \* Plane.point -> Plane.vec the vector whose tail is the first point and whose head is the second point
- Plane.zero : Plane.vec the zero vector
- Plane.add : Plane.vec \* Plane.vec -> Plane.vec
   add two vectors
- Plane.scale : Plane.vec \* Scalar.scalar -> Plane.vec scale a vector by a constant
- Plane.origin: Plane.point the origin point of the vector space.
- Plane.distance: Plane.point \* Plane.point -> Scalar.scalar Plane.distance(p1, p2) evaluates to the distance between the points p1 and p2.
- Plane.midpoint: Plane.point \* Plane.point -> Plane.point Plane.midpoint(p1, p2) evaluates to the midpoint of the points p1 and p2.

• Plane.head: Plane.vec -> Plane.point
Plane.head v evaluates to the point that corresponds to the displacement of v from the origin.

#### 2.1.2 Bounding boxes

The type BoundingBox.bbox represents a rectangular region in two-dimensional space. You will want to use the functions whose types and specs are given in bbox.sig. These functions are implemented in bbox.sml. Here some useful functions:

- BoundingBox.contained: (bool \* bool \* bool \* bool)
  \* Plane.point \* BoundingBox.bbox -> bool
  BoundingBox.contained(bs, p, bb) evaluates to true if and only if the point p is in the box b. The four booleans control whether the left/right/top/bottom edges of
  - in the box b. The four booleans control whether the left/right/top/bottom edges of the box are included or excluded, where true means to exclude an edge.

For example, BoundingBox.contains((false,false,false,false),p,b) returns true if p is in the box b including all of the edges, while BoundingBox.contains((true,false,false,false),p,b) is the same except it will return false if p is directly on the left edge of the rectangle. The order of the booleans is

(exclude left side, exclude right side, exclude top side, exclude bottom side)

and a corner is excluded if either of the sides containing it are excluded.

- BoundingBox.diameter: BoundingBox.bbox -> Scalar.scalar Computes the diameter of the box, i.e. the length of the diagonal.
- BoundingBox.from2Points: Plane.point \* Plane.point -> BoundingBox.bbox BoundingBox.from2Points (p1, p2) returns the smallest bounding box containing both p1 and p2
- BoundingBox.fromPoints: Plane.point Seq.seq -> BoundingBox.bbox Computes the minimum bounding box containing every point in a sequence of points, assuming the sequence is non-empty.
- BoundingBox.center : BoundingBox.bbox -> Plane.point Computes the center point of the bounding box.
- BoundingBox.corners : BoundingBox.bbox ->
  Plane.point \* Plane.point \* Plane.point \* Plane.point
  Returns the four corners of the bounding box in order

(top left, top right, bottom left, bottom right)

### 2.2 Barnes-Hut

In lecture, we discussed how to solve the *n*-body problem in the naïve, quadratic manner. The code for this is given in mechanics.sml and naiveNBody.sml. Recall that the pieces of information we need about a body in space are its mass, location, and velocity. This is represented by the type definition

```
type body = Plane.scalar * Plane.point * Plane.vec
```

The type body is used to represent the different bodies in the *n*-body simulation. Specifically, in an expression (m, p, v) of type body, m is the mass of the body, p is its position, and v is the vector representing its velocity.

The naïve, quadratic implementation of an n-body simulation is given by the function

```
accelerations : body Seq.seq -> Plane.vec Seq.seq
```

in naiveNBody.sml. This function transforms a sequence of bodies into a sequence in which the element at position i represents the acceleration for the element at position i of the sequence of bodies.

One of the vital helper functions for this is

```
accOn : body * body -> Plane.vec
```

found in mechanics.sml. Recall the specification is that accOn (b1, b2) calculates the acceleration on b1 due to b2. Using this function, the calculation is fairly straightforward:

```
fun accelerations (bodies : body Seq.seq) : Plane.vec Seq.seq =
    Seq.map (fn b1 => Plane.sum (fn b2 => accOn (b1, b2), bodies), bodies)
```

```
where Plane.sum(f,s) is Seq.reduce(Plane.add, Plane.zero, Seq.map(f,s))
```

However, on large inputs, this implementation is accurate, but unacceptably slow for an actual simulation. There are many different approximations that have been developed; the one we will look at is called Barnes-Hut.

#### 2.2.1 The algorithm

Barnes-Hut groups bodies by quadrants (in the two-dimensional case) and uses a threshold value  $\theta$  to determine whether each individual body is "far enough" away from a group of other bodies. If it is, it groups the other bodies into a big pseudobody and uses that for the acceleration calculation instead of each individual body composing the pseudobody. This results in a loss of accuracy, but a dramatic speedup in terms of runtime—while the old algorithm had work in  $O(n^2)$ , this algorithm's work is in  $O(n \log n)$  if the threshold value is well-chosen.

To calculate the effect of a pseudobody on another body, it is important to know the total mass of all the bodies represented by the pseudobody and also their center of mass or *barycenter*. Therefore, when we form a pseudobody, we will compute a tuple (m, c) such that m: Plane.scalar is the total mass of the bodies and c: Plane.point is the

barycenter. To compute the barycenter, we compute a weighted average of the vectors corresponding to the displacement of each body's position from the origin. For bodies  $(m_1, p_1, v_1), (m_2, p_2, v_2), \dots (m_n, p_n, v_n)$ , we compute the following vector:

$$\mathbf{R} = \frac{m_1 \mathbf{r}_1 + m_2 \mathbf{r}_2 + \dots m_n \mathbf{r}_n}{m_1 + m_2 + \dots m_n}$$

where  $\mathbf{r}_i$  is the vector from the origin to position  $p_i$ . The barycenter is then the head of this vector.

To approximate the acceleration on a body using a pseudobody, we represent the acceleration due to all the bodies in a group as the acceleration due to a single body located at the barycenter with mass equal to the total mass.

#### 2.2.2 Computing the barycenter

Rather than computing the barycenter for each quadrant of space from all of the points that space, we will compute the barycenter as an average of the averages of the four subquadrants of a region of space. This means that all we ever need to do is to compute the barycenter of four pairs of masses and points.

Task 2.1 (10 pts). Write the function

that computes the pair (m, c) where m is the total mass of the four bodies (i.e., the sum of the first components of the pairs) and c is the barycenter.

#### 2.2.3 Grouping bodies

We still have not discussed exactly how to group bodies. There are many different ways of doing so, but the most straightforward is by grouping things into quadrants (for the 2D case). That is, starting at the center of the area, we divide the field into quadrants, then recursively group the bodies in each quadrant, stopping when a region has either zero or one body in it. This yields a tree-structured division of space, where each node has four subtrees, corresponding to the four quadrants of it. We can represent this tree structure as a datatype in SML:

Empty represents a region with no bodies in it. Single b represents a region with exactly the body b in it. Cell ((m, c), bb, sq) is somewhat more complicated:

- m is the total mass of the bodies contained in the region.
- c is the barycenter of the bodies contained in the region.
- bb is a bounding box of the region.
- The four subtrees represent the subdivisions of the four quadrants of the region. The four child **bhtree**'s are, in order, the top-left, top-right, bottom-left, and bottom-right quadrants of the region, respectively.

As a first step in constructing this tree, we will write the quarters function to split a bounding box into four equally sized quadrants.

Task 2.2 (10 pts). Write the function

to compute the four bounding boxes that correspond to the top-left, top-right, bottom-left, and bottom-right quadrants of the argument bounding box. Use the bounding box functions described above.

#### 2.2.4 Growing the tree

We now have the tools we need to compute a **bhtree** from a sequence of points and a bounding box.

Task 2.3 (30 pts). Write the function

```
compute_tree : body Seq.seq * BoundingBox.bbox -> bhtree
```

such that compute\_tree (s, bb) evaluates to T, where T is the tree decomposition of s in the bounding box bb. You may assume that all of the bodies in s are within the bounding box bb and that no two bodies in s occupy the same position (i.e., have equal position components).

In the recursive calls, you will need to divide **s** into four sequences corresponding to those bodies in each of the four quadrants of **bb**. If a body is on the border between two quadrants, it should be placed in the *first* quadrant that it is in, in the following order: top left, top right, bottom left, bottom right.

Note: The barycenter of the bodies in a bounding box should be computed as the barycenter of the four quadrants' barycenters. You can use the helper function center\_of\_mass: bhtree -> Plane.scalar \* Plane.point to project the relevant data from the result of a recursive call.

#### 2.2.5 Computing acceleration

Now that we can calculate the tree determined by a group of bodies, we can use it to efficiently compute an approximation of the acceleration of all the bodies at this particular timestep. This brings us back to the threshold value  $\theta$  mentioned above.

The reason Barnes-Hut is more efficient than the naïve approach is that it does not compute the exact acceleration—instead, it uses a parameter  $\theta$  to determine exactly how precise to be. Whenever your algorithm reaches a region with more than one body in it (that is, a Cell in the tree), it checks to see if  $\frac{diam}{dist} \leq \theta$ , where diam is the diameter of the region (the length of the diagonal) and d is the distance from the body being checked to the region's barycenter. If it is, then the region is treated as one large body located at its barycenter (which we have conveniently already calculated!). Otherwise, the region gets decomposed into quadrants and the respective accelerations from the bodies in each quadrant are computed recursively, combined, and returned.

#### Task 2.4 (5 pts). Write the function

```
val too_far : Plane.point * Plane.point * BoundingBox.bbox * Scalar.scalar -> bool such that given a point p1 (position of body whose acceleration is being computed), a point c (center of a region), and bounding box bb (bounding box of that region) and a threshold t (\theta), too_far (p1, c, bb, t) evaluates to true if \frac{diam}{dist} \leq \theta and false otherwise.
```

#### Task 2.5 (25 pts). Write the function

```
bh_acceleration : bhtree * Plane.scalar * body -> vec
```

such that bh\_acceleration (T, threshold, b) computes the acceleration on b from the tree T according to the algorithm described above. (Hint: Use a function from mechanics.sig, which is the mechanics code from lecture.)

Task 2.6 (20 pts). Finally, write a function

```
barnes_hut : Scalar.scalar * body Seq.seq -> vec Seq.seq
```

that uses your compute\_tree and bh\_acceleration functions to form the Barnes-Hut tree for a sequence of bodies and then use it to compute the acceleration on each body in the sequence.

# 2.3 How to Load the Project

To load your code using the floating point implementation of the plane, in SMLNJ issue the command

```
- CM.make "sources.cm";
```

CM.make uses the SMLNJ compilation manager to load many different files, including all of our support code and the code that you write. Every time you edit your file, you should re-run CM.make to reload your code—do this in place of "using" the homework file.

### 2.4 Testing

Because your solution consists of several functions that build on each other, it is in your interest to test each function in isolation, which will help you figure out where your bugs are. For barycenter, quarters, and compute\_tree, we have provided some tests, and you can write more (if you want) using the helper functions and hard-coded points/bounding boxes/bodies in the module TestData. You can run BarnesHut.test\_X() for the test functions in BarnesHut.sml.

To test the overall behavior, you will generate transcripts describing the position of each body and look at them in a visualizer.

Use the following tests first:

• Transcripts.run\_solar\_inner : int \* string -> unit run\_solar\_inner (days, outfilename) generates a transcript for running the solar system for that many days and puts the results in the file data/outfilename. The transcript file tells the visualizer to show only the planets up to Mars.

For example:

```
- Transcripts.run_solar_inner (365, "year.txt.sim");
```

will create a file data/year.txt.sim that should show the earth (the blue circle) orbiting the sun once and the other planets orbiting as well.

• Transcripts.run\_solar : int \* string -> unit Same as the above, except the visualization radius includes all planets.

Once you have produced a transcript file, you can visualize it by navigating to

https://dlicata.wescreates.wesleyan.edu/teaching/fp-s23/visualizer/visualizer.html

You can then load a transcript file in one of two ways: either dragging and dropping the transcript file into the dashed box, or using the file browser to select the file manually. You should refresh the page before running another transcript.

Next, you should run some larger simulations. The support code contains initial conditions for the following:

# N-Body data files

galaxy10k.txt	10,000 bodies orbiting in an elliptical galaxy.
galaxy20k.txt	20,000 bodies orbiting in an elliptical galaxy.
galaxy30k.txt	Two 15,000 particle galaxies collide.
saturnrings.txt	Saturn with its 9 rings containing 11,987 particles.
cluster2.5k.txt	A star cluster containing 2,582 particles.
collision1.txt	Two large irregular galaxies collide and mostly pass
	through each other.
collision2.txt	Collision of two 1000 particle elliptical galaxies.
galaxyform2.5k.txt	A 2,500 particle dust cloud collapses under gravity

and forms into a galaxy. Two 1,000 particle clouds gravitate together and eventually galaxymerge1.txt combine into a galaxy. galaxymerge2.txt Four 1,000 particle clouds gravitate together and combine into a galaxy. A 900 particle spinning elliptical galaxy collides with a big galaxymerge3.txt expanding galaxy and combines with it. A lighter galaxy is torn apart by a dense galaxy. galaxy1.txt galaxy2.txt A galaxy orbits a heavier one and forms a spiral. galaxy3.txt A dense galaxy sucks stars out of a lighter galaxy that orbits it. galaxy4.txt Galaxies orbiting each other. A round galaxy becomes a spiral due to two small galaxies spiralgalaxy.txt that fly by it. Near earth asteroids and the earth in orbit around the sun. asteroids1000.txt

#### To run them:

• Transcripts.run\_file : string \* int \* real -> unit

run\_file (filename, num\_iters, timestep) runs the simulation on an input file specified by filename, for num\_iters steps, with time given by timestep. The output is written to filename.sim.

For example:

- Transcripts.run\_file ("data/galaxy2.txt", 2000, (1.0/10.0));

produces a file data/galaxy2.txt.sim. See data/datafiles.txt for descriptions of the simulations.

Here are some good timesteps and numbers of iterations.

```
Transcripts.run_file ("data/asteroids1000.txt", 1000, (1.0/10.0)); (* 62 seconds*)
Transcripts.run_file ("data/cluster2582.txt", 2000, (1.0/10.0)); (* 555 seconds *)
Transcripts.run_file ("data/galaxy1.txt", 2000, (1.0/10.0)); (* 130 seconds *)
\label{temperature} Transcripts.run\_file ("data/galaxy2.txt", 2000, (1.0/10.0)); (* 83 seconds *)
\label{temperature} Transcripts.run\_file ("data/galaxy3.txt", 1500, (1.0/10.0)); (* 304 seconds *)
Transcripts.run_file ("data/galaxy4.txt", 2000, (1.0/10.0)); (* 41 seconds *)
Transcripts.run_file ("data/spiralgalaxy.txt", 2000, (1.0/10.0)); (* 94 seconds *)
\label{temperature} Transcripts.run\_file ("data/galaxymerge1.txt", 5000, (1.0/5.0)); (* 820 seconds *)
Transcripts.run_file ("data/galaxymerge2.txt", 2500, (1.0/10.0)); (* 626 seconds *)
Transcripts.run_file ("data/galaxymerge3.txt", 2500, (1.0/10.0)); (* 655 seconds *)
Transcripts.run_file ("data/galaxyform2500.txt", 2000, (1.0/10.0)); (* 294 seconds
\label{tensor} Transcripts.run\_file ("data/collision2.txt", 2500, (1.0/10.0)); (* 330 seconds *)
Transcripts.run_file ("data/collision1.txt", 1500, (1.0/10.0)); (* 299 seconds *)
Transcripts.run_file ("data/saturnrings.txt", 100, (1.0/100.0)); (* 112 seconds *)
Transcripts.run_file ("data/galaxy10k.txt", 100, (1.0/10.0)); (* 93 seconds *)
Transcripts.run_file ("data/galaxy20k.txt", 50, (1.0/10.0)); (* 188 seconds *)
```

Transcripts.run\_file ("data/galaxy30k.txt", 800, (1.0/10.0)) (\* 1736 seconds \*)

Some of them take a while (the comment is how long they took for me). You don't need to run all of them — you can pick a few to try. Or you can either turn down the number of iterations to see less of the movie, or run them overnight. The total size of all files produced is about 2GB.

## • The command

- Transcripts.run\_files();

runs all of the above.